Spin-2 KK Mode Scattering in the Truncated RS1

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What am I presenting?

 <u>Research I performed with:</u> R. Sekhar Chivukula, Kirtimaan A. Mohan, Dipan Sengupta, and Elizabeth H. Simmons

Our Relevant Papers:

- "Sum Rules for Massive Spin-2 Kaluza-Klein Elastic Scattering Amplitudes" [arXiv:1910.06159], Phys. Rev. D 100, 115033
- "Scattering Amplitudes of Massive Spin-2 Kaluza-Klein States Grow Only as O(s)" [arXiv:1906.11098]

 "Scattering Amplitudes and Strong Coupling Scales in Extra-Dimensional Theories" [arXiv: 2002.XXXX]

** COMING SOON! **

The Overall Idea

4D Spacetime \rightarrow 4D Graviton



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4D Spacetime \rightarrow 4D Graviton





$$g_{\mu\nu} = \eta_{\mu\nu} + \underbrace{\frac{2}{M_{\rm Pl}} \hat{h}_{\mu\nu}^{(0)}}_{\text{disturbance}}$$
$$\mathbf{M}_{\rm Pl} = \text{the reduced Planck mass}$$

 $ds^2 = g_{\mu\nu} \, dx^\mu \, dx^\nu$

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4D Spacetime \rightarrow 4D Graviton





Graviton

 $\hat{h}^{(0)}_{\mu\nu} = 4D$ massless spin-2 field \equiv the "4D graviton" field



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2-to-2 Scattering in 4DG





2-to-2 Scattering in 4DG





2-to-2 Scattering in 4DG







2⁴ = 16 helicity combos

Because...

- **M**_{Pl} is the only dimensionful parameter.
- The 4D matrix element is dimensionless.

2-to-2 Scattering in FPG



Fierz Pauli massive gravity (FPG) is constructed by adding a **Fierz-Pauli mass term** to 4D gravity, which explicitly breaks 4D diffeomorphism invariance.

The field h⁽ⁿ⁾ describes a **massive** spin-2 particle.



 $5^4 = 625$ helicity combos

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Example: 5D Spacetime \rightarrow 5D Graviton



$$\eta_{MN} = \left(\begin{array}{ccccc} 0 & - & 0 & 0 & 0 \\ 0 & 0 & - & 0 & 0 \\ 0 & 0 & 0 & - & 0 \\ 0 & 0 & 0 & 0 & - \end{array}\right)$$

Suppose new extra dimension has finite length πr_c where r_c = the compactification radius

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Example: 5D Spacetime \rightarrow 5D Graviton





$$G_{MN} = \eta_{MN} + \underbrace{\kappa \hat{h}_{MN}}_{\text{disturbance}}$$

$$\kappa = \text{the 5D coupling strength}$$

and has units [E]^{-3/2}

$$ds^2 = G_{MN} \, dx^M \, dx^N$$

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Example: 5D Spacetime \rightarrow 5D Graviton







```
\hat{h}_{MN} \equiv the "5D graviton" field
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Kaluza-Klein Decomposition: Hope for Massive Spin-2?



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The "Punchlines" Slide



We examine how this constraint impacts the
 5D Randall-Sundrum (RS1) model, which
 possesses warping in the extra dimension.

- We perform the full calculation, i.e. we don't rely on the "large warping" limit.
- We derive sum rules relating couplings & masses between KK modes.
- Cancellations of E¹⁰ behavior require infinitely many modes. When truncating the KK sum, many states may be needed to sufficiently approximate the full matrix element.

The Randall-Sundrum "RS1" Model

The Original RS Papers: [arXiv:hep-ph/9905221 and 9906064]

Previous Slide

N/A







Add an extra spatial dimension with coordinate **y** ranging from $\mathbf{y} = \mathbf{0}$ to $\mathbf{y} = \pi \mathbf{r}_c$ where \mathbf{r}_c is called the compactification radius.





The compactification radius r_c proves useful for creating dimensionless quantities. For instance, we'll sometimes replace **y** with $\phi = y/r_c \in [0,\pi]$

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The warping parameter k (dimensions of energy) measures spacetime distortion.

The case $\mathbf{k} = \mathbf{0}$ describes a flat extra dimension.

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Orbifold

Extend spacetime by reflecting across $\varphi = 0$ so that $\varphi \in [-\pi, +\pi]$.

By demanding that the interval **ds**² is invariant under this reflection, we construct an **orbifolded** theory.

This projects onto **orbifold even states** (including the **graviton** + **radion**) and eliminates orbifold odd states (graviphoton)

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Previous Slide



The RS1 Metric & Field Content

$$\eta_{MN}^{(\mathrm{RS})} \xrightarrow{\mathrm{perturb}} G_{MN}^{(\mathrm{RS})} = \begin{pmatrix} \varepsilon^{-2} g_{\mu\nu} e^{-2\hat{u}} & 0\\ 0 & -(1+2\hat{u})^2 \end{pmatrix}$$

where

$$g_{\mu\nu} \equiv \eta_{\mu\nu} + \kappa \hat{h}_{\mu\nu}$$
 and

$$\hat{u} \equiv \frac{\kappa}{2\sqrt{6}} \left[\varepsilon^{+2} e^{-\pi k r_c} \hat{r} \right]$$

 $\hat{h}_{\mu\nu}(x,y) = \text{massless 5D spin-2}$

 $\hat{r}(x) = \text{massless 5D spin-0} \quad (\text{note it's y-independent!})$

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Building RS1: Kaluza-Klein (KK) Decomposition



$$\hat{h}_{\mu\nu}(x,y) = \frac{1}{\sqrt{\pi r_c}} \sum_{n=0}^{+\infty} \hat{h}_{\mu\nu}^{(n)}(x) \, \underline{\psi}_n(\varphi)$$

= (4D graviton) + (massive 4D spin-2 KK modes)

$$\underline{\hat{r}(x)} = \frac{1}{\sqrt{\pi r_c}} \underline{\hat{r}^{(0)}(x)} \underline{\psi_0}$$

 The wavefunctions ψ_n solve a specific Sturm-Liouville Equation (S-L Eq.) The solution set { ψ_n } and spectrum
 { m_nr_c } are entirely determined by kr_c.

Building RS1: The Lagrangian

$$\begin{array}{l} \text{RS1 Metric} \\ G_{MN}^{(\text{RS})} = \left(\begin{array}{cc} \varepsilon^{-2} g_{\mu\nu} e^{-2\hat{u}} & 0 \\ 0 & -(1+2\hat{u})^2 \end{array} \right) \\ \text{where} \quad \varepsilon \equiv e^{+kr_c|\varphi|} \end{array}$$

$$\mathcal{L}_{5D} = \mathcal{L}_{EH} + \mathcal{L}_{CC} + \Delta \mathcal{L}$$

$$\mathcal{L}_{\rm EH} = \frac{2}{\kappa^2} \sqrt{\det G} \ R_{\rm 5D}$$

Cosmological Constant Terms

$$\mathcal{L}_{\rm CC} = \frac{12}{\kappa^2} k r_c \left\{ 2\sqrt{\det G} \left(\partial_{\varphi} |\varphi|\right)^2 - \varepsilon^{-4} e^{-4\hat{u}} \sqrt{-\det g} \left(\partial_{\varphi}^2 |\varphi|\right) \right\}$$

to ensure 4D effective theory is flat.

• Total Derivative

$$\Delta \mathcal{L} = \frac{2}{\kappa^2} \partial_y \left[\frac{\varepsilon^{-4} e^{-4\hat{u}}}{\sqrt{1+2\hat{u}}} \sqrt{-\det g} \left(g_{\mu\nu}^{-1} (\partial_y g^{\mu\nu}) + \partial_y \ln \left[\varepsilon^{-4} e^{-4\hat{u}} \right] \right) \right]$$

to eliminate all instances of ∂_v^2 for later convenience.

Reparameterize: $(k, r_c, \kappa) \rightarrow (kr_c, m_1, M_{Pl})$

$$kr_{c} = k \cdot r_{c}$$

$$m_{1} \equiv \frac{1}{r_{c}} \left[(m_{1}r_{c}) \big|_{kr_{c}} \right] \text{ via S-L Eq.}$$

$$M_{\text{Pl}} \equiv \frac{2}{\kappa\sqrt{k}} \sqrt{1 - e^{-2kr_{c}\pi}}$$

For the remainder of this presentation, whenever I show *numerical* plots, I set:

$$M_{\rm Pl} = 2.435 \times 10^{15} \text{ TeV}$$
$$m_1 = 1 \text{ TeV}$$

$$\mathcal{L}_{5\mathrm{D}} = \mathcal{L}_{\mathrm{EH}} + \mathcal{L}_{\mathrm{CC}} + \Delta \mathcal{L} \qquad \stackrel{\mathrm{WFE}}{\underset{\mathbf{\& \ KK}}{\longrightarrow}} \qquad \mathcal{L}_{4\mathrm{D}}^{(\mathrm{eff})} \equiv \int dy \quad \mathcal{L}_{5\mathrm{D}}$$

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Example: Cubic KK Mode Interaction



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Example: Quartic KK Mode Interaction



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Example: (KK Mode)² - Radion Interaction



If two extra-dimensional derivatives...

B-Type Coupling

No A-Type (KK Mode)² - Radion Interaction Appears in L_{5D}

$$\varepsilon^{-4} \left[\varepsilon^{+2} e^{-\pi k r_c} \hat{r} \right] (\partial_{\varphi} \hat{h}) (\partial_{\varphi} \hat{h}) \longrightarrow \varepsilon^{-2} (\partial_{\varphi} \psi_{n_1}) (\partial_{\varphi} \psi_{n_2}) \psi_0$$

$$\longrightarrow \left| b_{n_1 n_2 r} \equiv \frac{\psi_0}{\pi r_c} e^{-\pi k r_c} \int d\varphi \ \varepsilon^{-2} \left(\partial_{\varphi} \psi_{n_1} \right) \left(\partial_{\varphi} \psi_{n_2} \right) \right|$$

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Matrix Element: Total Matrix Element



$$\mathcal{M}_{\text{full}} = \frac{\lambda_1 \mathbf{n}_1 \mathbf{n}_2 \mathbf{n}_3 \lambda_3}{\lambda_2 \mathbf{n}_2 \mathbf{n}_2 \mathbf{n}_4 \lambda_4} = \mathcal{M}_c + \mathcal{M}_r + \sum_{j=0}^{+\infty} \mathcal{M}_j$$

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Matrix Element: Truncated Matrix Element



 $\mathcal{M}^{[N]} \equiv \mathcal{M}_c + \mathcal{M}_r + \sum^{\mathbf{N}} \mathcal{M}_j$

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Matrix Element: Expanded Matrix Element



Expectation: $\underline{\mathcal{M}}^{[N](\sigma)} \equiv \sum \ \overline{\mathcal{M}}^{[N](\sigma)} \ (Er_c)^{\underline{2\sigma}}$ $M^{[N](\sigma)}$ vanishes as $N \rightarrow \infty$ for all $\sigma > 1$ σ σ (i.e. growth > E^2)

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Cancelling Bad High-Energy Growth Sum Rules & Numerical Checks

Longitudinal $(1,1) \rightarrow (1,1)$: Cancellations



Q: How does including more terms in the truncated sum affect leading energy growth?

A: By increasing N, any energy growth faster than O(E²) increasingly cancels. O(E²) remains.



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Rewrite Generic Elastic Couplings



How does this happen analytically?



Focus on the **longitudinal elastic scattering matrix elements**, which seemingly grow like **O(E**¹⁰) at finite truncation.

Use wavefxn properties to rewrite **B-Type couplings** as **A-Type couplings**:

$$b_{nnj} = \left[(m_n r_c)^2 - \frac{1}{2} (m_j r_c)^2 \right] a_{nnj}$$

$$b_{jnn} = \frac{1}{2} (m_j r_c)^2 a_{nnj}$$

$$b_{nnnn} = \frac{1}{3} (m_n r_c)^2 a_{nnnn}$$

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Sum Rules (order-by-order)

 $\pm \infty$

The **longitudinal elastic scattering matrix elements** grows like O(E²) iff the following **sum rules** between masses & couplings hold true:

$$\begin{split} \underline{\mathcal{O}(s^5)} &: \qquad \sum_{j=0}^{+\infty} a_{nnj}^2 = a_{nnnn} \\ \underline{\mathcal{O}(s^4)} &: \qquad \sum_{j=0}^{+\infty} \left[\frac{m_j}{m_n}\right]^2 a_{nnj}^2 = \frac{4}{3} a_{nnnn} \\ \\ \underline{\mathcal{O}(s^3)} &: \qquad \sum_{j=0}^{+\infty} \left[\frac{m_j}{m_n}\right]^4 a_{nnj}^2 = \frac{4}{5} \left[9 \frac{b_{nnr}^2}{(m_n r_c)^4} - a_{nn0}^2\right] + \frac{16}{15} a_{nnnn} \\ \\ \underline{\mathcal{O}(s^2)} &: \qquad \sum_{j=0}^{+\infty} \left[\frac{m_j}{m_n}\right]^6 a_{nnj}^2 = 4 \left[9 \frac{b_{nnr}^2}{(m_n r_c)^4} - a_{nn0}^2\right] \end{split}$$

Reorganize These Last 2

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Sum Rules (alternate set)

The **longitudinal elastic scattering matrix elements** grows like O(E²) iff the following **sum rules** between masses & couplings hold true:

$$\begin{split} \underline{\mathcal{O}(s^5)} &: \quad \sum_{j=0}^{+\infty} a_{nnj}^2 = a_{nnnn} \quad \textbf{PROVED} \\ \underline{\mathcal{O}(s^4)} &: \quad \sum_{j=0}^{+\infty} \left[\frac{m_j}{m_n} \right]^2 a_{nnj}^2 = \frac{4}{3} a_{nnnn} \quad \textbf{PROVED} \\ \underline{(s^2)} \leftrightarrow \mathcal{O}(s^3) &: \quad \sum_{j=0}^{+\infty} \left[\left(\frac{m_j}{m_n} \right)^2 - 5 \right] \left[\frac{m_j}{m_n} \right]^4 a_{nnj}^2 = -\frac{16}{3} a_{nnnn} \quad \textbf{PROVED} \\ \underline{\mathcal{O}(s^2)} &: \quad \sum_{j=0}^{+\infty} \left[\frac{m_j}{m_n} \right]^6 a_{nnj}^2 = 4 \left[9 \frac{b_{nnr}^2}{(m_n r_c)^4} - a_{nn0}^2 \right] \quad \begin{array}{c} \text{This is the only} \\ \text{rule left without} \\ \text{an analytic proof} \end{array}$$

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Elastic 2-to-2 KK Mode Scattering Matrix Elements in RS1 Fastest Energy Growth per Helicity Combination: $(\lambda_1, \lambda_2) \rightarrow (\lambda_3, \lambda_4)$



all

n



40 of 46

N n

all

Longitudinal $(1,4) \rightarrow (2,3)$: Cancellations



Q: How does including more terms in the truncated sum affect leading energy growth?

A: By increasing N, any energy growth faster than O(E²) increasingly cancels. O(E²) remains.



Finite Truncation at Low Energy How many states should I include?

Truncation at Low Energies: $E = 10 m_1$



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Truncation at Low Energies: $E = 10 m_1$



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Truncation at Low Energies: $E = 100 m_1$





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Conclusion

- Interested in learning more? These results (and more!) will be available on the arXiv this month.
 - See also: [arXiv:1906.11098] ; [arXiv:1910.06159]
- <u>What's next?</u>
 radion stabilization
 dark matter applications
 high-energy behavior of RS1 bulk/brane scalar matter



Thank you for attending!